

Electron Impact Ionication in Semiconductors.

Jan Tauc

Institute of Technical Physics, Czechoslovak Academy of Sciences, Frague, Guechoslovakia.

#### Abstract.

Experimental facts are presented showing that the quantum yield of the inner photoelectric effect increases if the photon energy exceeds a certain value. This is interpreted as due to the intrinsic electron impact ionisation. A theory of this effect is discussed. The relation of the high energy quantum yield to the low energy parameters is shown. Some other effects related to the intrinsic impact ionisation are mentioned.

If the kinetic energy of an electron or a hole attains a certain value an interaction with the valence band electrons can take place leading to the generation of a new electron-hole pair. In semiconductors this intrinsic impact ionisation was first observed in the breakdown of silicon (1-5) and germanium (6-8) pen junctions. Considering the breakdown as a solid state smalog of the Townsend  $\beta$  -avalanche breakdown in gases it was possible to deduce some fundamental information on the impact ionisation processe.

Approved For Release 2009/07/09: CIA-RDP80T00246A004400310001-8

# 1. Experimental Studies of the Quantum Yield.

We have dealt with this effect from another standpoint studying the dependence of the quantum yield of the inner photoelectric effect on the energy of the absorbed photon. The quantum yield  $\eta$  is the number of the electron-hale pairs created on the absorption of one photon. It was known from the measurement with low energy photons that the quantum yield near the infrared limit is one. From the measurement with high energy photons in the region of X - rays it followed that m is propertional to the energy of the photon (9) or, in other words, in this region the energy needed for the creation of one electron-hole pair is constant (in Ge  $\mathcal{E}$  = 2.5 eV). It was obviously of interest to study the transition between the two regions. The measurement was performed by Koe (10) (Fig.1). Similarly as in the work (9) he used the photovoltaic effect in a germanium p-n junction. It is apparent that the quantum yield  $\eta$  increases with the photon energy  $z_{\nu}$  if it exceeds the value E' = 2.2 eV. The tangent at the point B, = E, passes through the origin. It is remarkable that its slope is nearly the same as the slope of the corresponding straight line =  $\eta$  (E, ) in the X-ray region. A more detailed examination of the curve reveals a certain structure the nature of which can be shown more clearly in InSb (see below).

The ressurement of  $\eta$  in Ge was recently repeated by Vavilov and Britain (11) by a similar method. They have found qualitati-voly the care curve but the critical energy E, was found to be 2.9 eV. The Character he probably due to different methods of recommend the southern who sould have and Britain measured the resulted vity using a hollow sphere coated with 600 thus taking

Approved For Release 2009/07/09: CIA-RDP80T00246A004400310001-8

terty agreemed different light which was maglected by Nec.

Recently, Variety and Dritain have measured the quantum yield in allieum by a similar method; the energy  $H_{\nu}^{\prime}$  was found to be about 5.25 eV  $^{(12)}$ .

The uncertainty in the reflectivity is much smaller with InSt where it is practically constant near the energy E, . With this meterial it is difficult to use the pen Americas but it is easy to measure the photomogneteelectric and the photoconductive effects. The fermer effect increases with increasing absorption exefficient, the latter degreeses. If the two effects have the same dependence on E. it is considered probable that the change of the absorption coefficient with E, has negligible influence on the result. This was the case in our measurements (15), a twpical result of which is in Fig. 2. The increasing of 9 with E. starte at E. = 0.47 eV at room temperature. The tengent at this point passes again through the origin but the rise in the curve seen diminishes and even stops, At the energy R," = 0.6 eV starts to ingress again with a lover slope than at the point I' . Although the exact shape of the curve was not well reproduedble with different samples and surface treatments nevertheless the points E', E" were always noticeable. They nove towards higher values with degressing temperature; the order of magnitude of this change is 10<sup>-4</sup> eV/grad.

### 2. Errediction with Meetrons.

The quantum yield was also determined in the high energy region with irradiation by fact electrons (14-16) and  $\alpha$  -particles (17,18). These measurements gave values of  $\alpha$  somewhat higher

which with with proper we have tried to measure  $\gamma$  in germentum with low energy electrone. For (19) used the electron-voltate effect in germentum p-s imetions (grown or alloyed). The measurements showed the existence of a certain threshold energy  $\mathbf{E}_{\min}$ , below which no signal was observed. For the energy of the electrone  $\mathbf{E} > \mathbf{E}_{\min}$   $\gamma$  is a linear function of  $\mathbf{E}$ . The threshold energy  $\mathbf{E}_{\min}$  is several hundred volts and changes with the condition of the surface.

We suppose that Emin corresponds to the energy necessary for the electrons to penetrate through the oxide layer on the surface. The estimate of the thickness of the layer based on this assumption agrees with that obtained by other methods (tens of Angetrius).

Threnberg et al. (20) have found an energy threshhold of several thousands of eV in selenium barrier layer cells. Li Tschi-Tsmian has bombarded GSS and GSSs polycrystalline layers with very slew electrons and measured the changes of the conductivity (21). He has found a very low threshhold of the incident energy (tenths of an electronwolt) and abrupt jumps in the signal at the multiples of the energy gap (2.5 eV in GSS, 1.8 eV in GSSs).

### 3. Interpretation of the Low Energy Ourve.

by the following mechanism. On the absorption of a photon an electron in the conduction band and a hole in the valence band are created which have in the first moment excess kinetic energies. If these energies are lower that the ionisation energies necessary for the creation of a new electron-hole pair they are

transferred to the lattice and lost. If the kinetic energy of an electron or a hole exceeds the ionisation energy new pairs are created by an impact ionisation and the quantum yield is larger than one. The number of the created electron-hole pairs increases with the excess energies of the electron E<sub>n</sub> or the hole E<sub>p</sub> .

A theory based on this idea was worked out by Antoncik (22). He considers the ionising action only the electrons in the conduction band. The minimum chergy Enton which an electron must have to be able to ionise the valence band electrons is determined by the conservation laws for energy and crystal momentum in the process

## 1 fast electron -> 1 hole + 2 electrons

In, ion depends on the band structure. With a somewhat schematised model of the band structure Antonöik obtained for Ge in, ion = 0.82 eV = 1.24  $E_0$ , for Si 1.25 eV = 1.14  $E_0$  (22).

Antoncik then colsplates the number of electron-hole pairs created by impact ionisation in the neighbourhood of  $E_n$ , ion; assuming that a fast electron can ionise only once. If his bend structure is not spherically symmetrical (as is the case for Ge, Si, InSb) the conservation laws are fulfilled for  $E_n = E_n$ , ion only in a finite number of directions in the Brillouin zone. In Ge these are the (111) directions connecting the centre of the zone with the minima of the conduction band. Antoncik then show that if  $E_n > E_n$ , ion the conservation conditions hold in a certain solid angle around the starting (111) direction which inspecases with  $E_n$ . Supposing that all electrons in this solid angle give now pairs by ionisation, one obtains for the number

of the pairs created by impact ionisation (corresponding to one absorbed photon);

$$\eta_{in} = \frac{1}{2\sigma} \frac{1}{2\sigma} \frac{1}{2\sigma} \left( \mathcal{R}_{i} - \mathcal{R}_{n, 1cm} \right) \tag{2}$$

This formula helds for schematised many valley conduction bands such a these of Ge and Si, a is the number of minima in the conduction band, the number of ellipsoids ( $\mathcal{F} = \frac{1}{2} + 8$  for Ge,  $\mathcal{F} = 8$  for Si).  $\infty$  and D are parameters depending on the band structure ( $\mathcal{K} = 0.321$ , D = 1.242 for Si,  $\mathcal{K} = 0.305$ , D=1.525 for Ge (22)).

The ionisation actually starts at some energy  $E_n > E_n$ , ion owing to the interaction of the electron with the lattice. As the cross-section for the impact ionisation very steeply increases with the electron energy  $E_n$  (once it exceeds  $E_{n,ion}$ ) whilst the cross-section for the interaction with the lattice electry varies with  $E_n$ , the energy  $E_n$  should not differ such from  $E_{n,ion}$  (some tenths of an ev (22)).

In principle, similar considerations should held for the impact ionisation by heles.

if we are going to apply these considerations for the determination of m the question arises how are the energies of electrons and holes distributed just after the absorption of a photom. This is a difficult problem not solved yet. Antoncik (22) makes a very rough assumption that in Gs the electron takes all the excess energy. Using Eq (1) one obtains for the quantum yield for E, slightly larger than E,

where  $\Delta = 0$  (1 +  $\propto$ )/26 DE<sub>n,ion</sub> = 1 eV<sup>-1</sup>. This is in reasonable agreement with experiment. The energy E<sub>n</sub> = E<sub>V</sub> - E<sub>U</sub> 1.5 eV is considerably higher than E<sub>n,ion</sub> = 0.82 eV and agreement with the value determined from the breakdown of the p-n junctions by Miller <sup>(6)</sup>. Using the same argument for Si we get with the data from Vavilov and Britsin <sup>(12)</sup> E<sub>n</sub> = 3.25 - 1.1 = 2.25 eV which is in better agreement with the result obtained by Chynoweth and Me Eay <sup>(5)</sup> (2.25 eV) than with that by Miller <sup>(4)</sup> (1.5 eV).

The assumption that the electron has all the excess energy is certainly not well fulfilled in Ge and Si and we should expect that the actual values of  $E_n$  are somewhat lower. But it may be a very good assumption for InSb when the effective mass of an electron is far smaller than that of a hole. The observed curve  $\eta = \eta$  (E, ) in Fig. 2 could be explained by a band structure echematically suggested in Fig. 3. For small energy differences  $E_n = E_0$  practically all the energy  $E_n = E_0$  is taken by the electron. With increasing  $E_n = E_0$  the holes obtain a larger and larger part of this energy. This energy is lost as long as the holes do not attain the ionisation energy. Therefore in this region the initial rise of  $\eta$  with  $E_n$  slows down and eventually stops. As soon as the holes are able to ionise  $\eta$  starts to increase again (at the point  $E_n^{(r)}$ ).

An essential point in the theory of Antoncik is the anisotropy of the band structure. For a spherically symmetrical band structure 7 should change abruptly at E, (similarly as was observed in ref. (21)).

## 4. High Energy Region.

Where increasing the photon energy we get into a region where at present there is neither experimental nor theoretical information available. The situation is simple again at high energies  $E_{\gamma} >> E_{q}$  where  $\gamma$  is proportional to the energy of the photon or the particle. We may assume that in any case high energy electrons are first created which loose their energy by impact ionisation. The process is unddubtedly complex and complicated by emissions and reabsorptions of photons. However if we take the energy losses by secondary X-ray emission into account we find that the not energy  $\mathcal E$  needed for the creation of an electron-hole pair is little dependent on the energy of the photon (9). It follows that electrons with high energies  $(E_{n} \gg E_{n}')$  transfer little energy to the lattice. This is also to be expected on simple theoretical considerations.

By successive ignisation processes the energies of the electrons and holes finally sink under  $E_n$ ,  $E_p$ . If we suppose that no losses to the lattice occur for  $E_n > E_n$ ,  $E_p > E_p$  we have

$$\mathcal{E} = \mathbb{E}_{\mathbf{G}} + \langle \mathbb{E}_{\mathbf{p}} \rangle + \langle \mathbb{E}_{\mathbf{p}} \rangle \tag{5}$$

where  $\langle E_n \rangle$ ,  $\langle E_p \rangle$  are the mean values in the intervals  $0, E_n$  and  $0, E_p$ . Supposing, as is occasionally done  $^{(4, 5)}$ , that the energy is uniformly distributed in these intervals we have  $\mathcal{E} = E_0 + \frac{1}{2} E_n + \frac{1}{2} E_p$ .

The value determined by Eq. (3) must be considered as the lowest limit of  $\epsilon$ . We know from the measurement of  $\eta$  that not all electrons and holes with energies somewhat greater than

ω ?) **ω** 

En, Ep can ionise. A part of the energy in this region is necessarily lost to the lattice. Using here the experimentally determined parameters we can calculate & supposing we know the distribution of the energies of electrons and heles. This is a rather difficult problem not yet satisfactorily solved. The relation between the high and low energy effects are being experimentally and theoretically studied in our laboratory.

### 5. Other Effects Related to Impact Ionisation.

Most of the information on the ionisation process during the breakdown of p-n junctions was obtained by an analysis of the electrical data. However, it is to be expected that interesting facts, especially on the energy distribution of electrons and holes, may be deduced from the study of electroluminescence (23,24) and external electron emission (25, 26). Important deductions were obtained by the study of the yield of the external photoelectric emission that appreciably diminishes if the energy of the excited electrons attain the ionisation energy (27).

Recently Landsberg has been able to predict the rise of the quantum yield with Ey in InSb from a comparison of the radiative lifetimes determined by direct measurement of the radiation or by calculation from the optical constants (28).

Thanks are due to Dr. E. Antoncik for his helpful discussions during the preparation of this paper.

#### Taferences.

77

V	
ı	K.G. Mc Kay, K.B. Mc Afee, Phys. Rev. 91, 1079 (1955)
2	K.G. Mc Kay, Phys. Rev. 94, 877 (1954)
3	P.A. Wolff, Phys. Rev. 95, 1415 (1954)
4	S.L. Miller, Phys. Rev. 105, 1246 (1957)
5	A.G. Chynoweth, K.G. Me Kay 108, 29 (1957)
6	S.L. Miller, Phys. Rev. 99, 1234 (1955)
7	B.M. Vul, Zh.techn.fiz. 26, 2403 (1956)
8	A.P. Shotov, Zh.exp.theor.fiz. 26, 1634 (1956);
	Zh.techn.fiz. 28, 437 (1958)
9	J. Drahokoupil, W. Malkovská, J. Tauc, Gzechosl. J. Phys.
	7, 57 (1957)
10	S.Koc, Czechosl. J. Phys. 7, 91 (1957)
11.	V.S. Vavilov, K.I. Britsin, Zh.exp.theor.fiz. 34, 521 (1958)
12	V.J. Vavilov, K.I. Britain, Zh.exp.theor.fiz. 34, 1354 (1958)
13	J. Tauc, A. Abrahém, Ozechosl. J. Phys. (to be published)
14	P. Rappaport, Phys. Rev. 93, 246 (1954)
15	W.G. Pfann, W.van Rocebroeck, J.Ropl. Physics 25, 1422 (1954)
16	V.S. Vavilov, L.S. Smirnov, V.M. Packevitsh, Dokl. Aked.
Ar Ar	Nauk 112, 1020 (1957)

Approved For Release 2009/07/09 : CIA-RDP80T00246A004400310001-8

K.G. Mc Key, Phys. Rev. 84, 829 (1951)

28

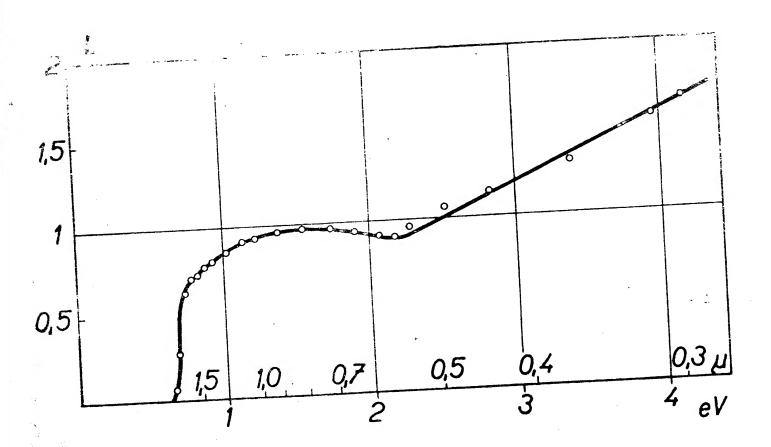
K.G. Mc Kay, K.B. Mc Affee, Phys.Rev. 91, 1079 (1995) 18 19 S. Koe (to be published) 50 W. Ehrenberg, Chi-Shi Long, R. West, Proc. Phys. Soc. (London) A 64, 424 (1951) Li Chi-Tszian, Dissertation, University, Leningred, 1958 21 E. Antončík, Cmechosl. J. Phys. 7, 674 (1957) 22 23 R. Newman, Phys. Rev. 100, 700 (1955) A.G. Chynoweth, K.G. Mc Kay, Phys. Rev. 102, 369 (1956) 24 J.A. Burton, Phys.Rev. 108, 1342 (1957) 25 J. Tauc. Nature 181, 38 (1958) 26 27 E.A. Taft, H.R. Philipp, L. Apker (to be published)

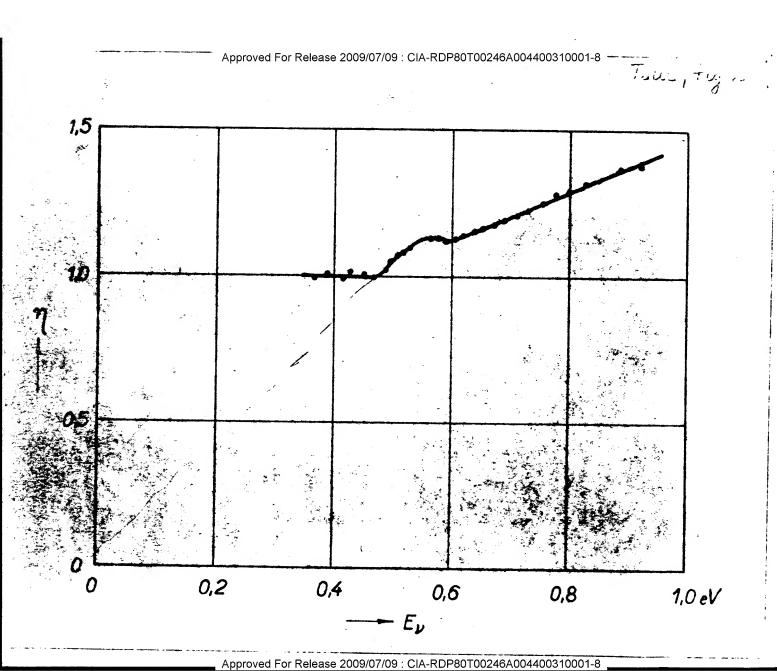
P.T. Landsberg, Proc. Phys. Sec. (London) B 70, 1175 (1957).

### Text to the Figures.

- Fig. 1 Quantum yield 7 in Ge as a function of the energy of the photon B,
- Fig. 2 Quantum yield m in Inst
- Pig. 3 Suggested band structure of InSb to explain qualitatively the curve given in Fig. 2.







Approved For Release 2009/07/09: CIA-RDP80T00246A004400310001-8

Tauc, Fig. 3

